

BREAKTHROUGH LISTEN

DANNY PRICE
UNIVERSITY OF CALIFORNIA BERKELEY
HDF WORKSHOP, 19 JUNE 2016

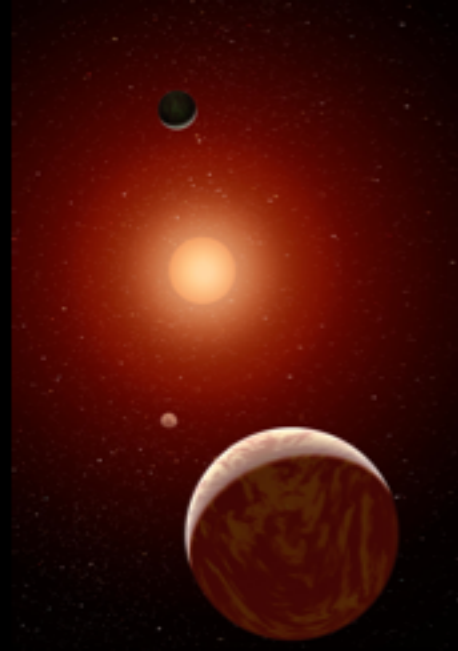


BERKELEY SETI
RESEARCH CENTER

Breakthrough Listen is the largest ever scientific research program aimed at finding evidence of civilizations beyond Earth.

THE BREAKTHROUGH LISTEN INITIATIVE:

OVERVIEW



1 Million Stars



*Milky Way Galactic Plane
Survey*



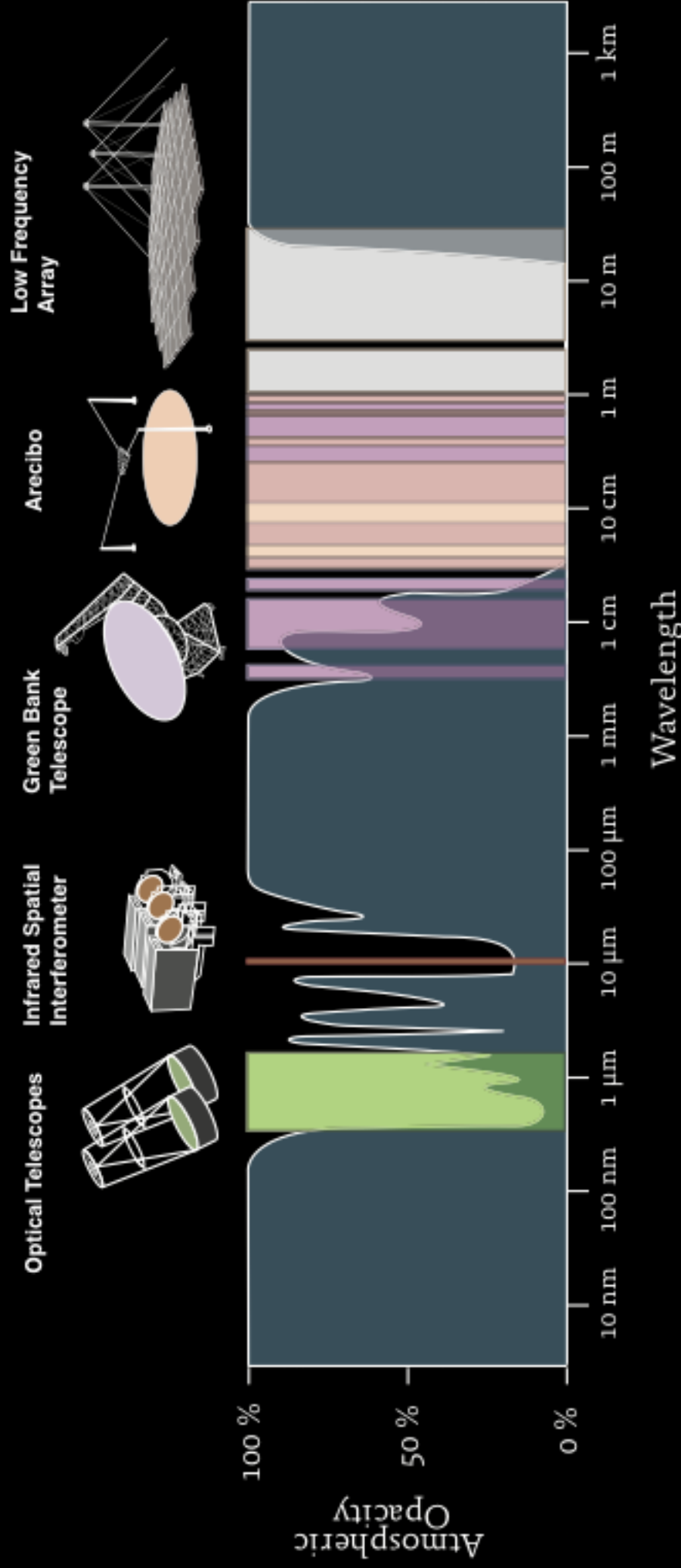
100 Galaxies

1 day of Breakthrough Listen = 1 year of any previous search

Many PB of scientific data to be made publicly available:

BL will be one of the largest public astronomy datasets

SEARCHING ACROSS THE ELECTROMAGNETIC SPECTRUM



Using multiple telescopes, we can search across the electromagnetic spectrum for indicators of advanced technology.

THE BREAKTHROUGH LISTEN INITIATIVE: TELESCOPES



Automated Planet Finder (Lick Observatory)

- Search for extremely narrow emission lines from artificial lasers
- Extremely high resolution “Levy Spectrometer”
 $374 - 950 \text{ nm}, \lambda/\Delta\lambda = 10^5$



Green Bank Telescope (Green Bank, WV)

- Radio search focusing on targeted and raster observations
- Nearly continuous frequency coverage 300 MHz - 100 GHz
- Flexible IF system can deliver up to 10 GHz dual-pol analog bandwidth
- Extremely radio quiet (Federally protected radio quiet zone)

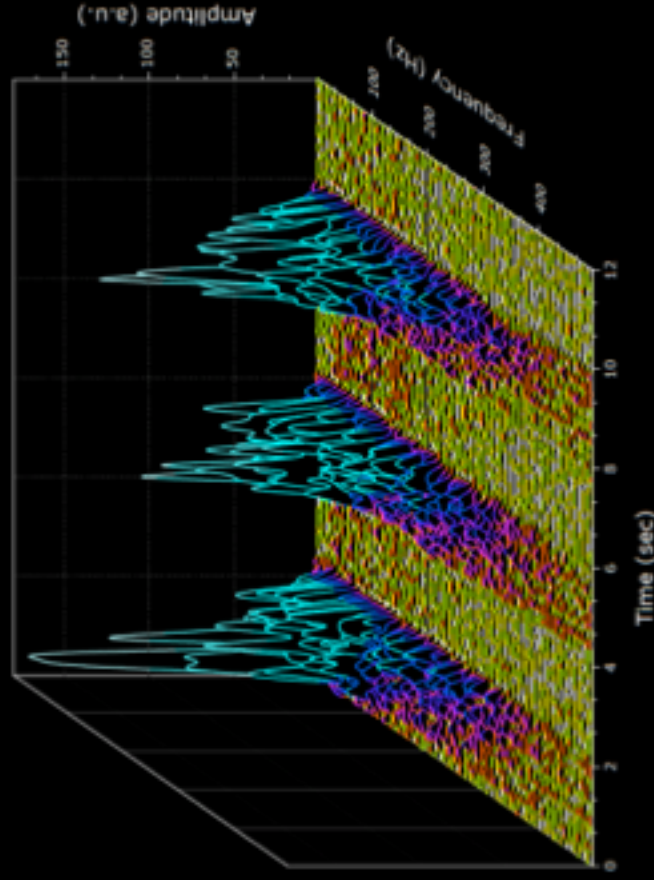


Parkes Telescope (NSW, Australia)

- Radio search focusing on surveys
- Southern hemisphere location gives great access to galactic plane
- Multi-beam receiver allows very efficient L-band (1.2 - 1.5 GHz) galactic plane surveys (Parkes Multibeam Receiver)
- Wide-band single-pixel and Phased Array Feed upgrades possible.

DATA OVERVIEW

- Data are N -dimensional arrays of power spectra (PSD)
- Axes are (frequency, time) and each data point is power
- Telescope metadata also needs to be stored (where it's pointing!)
- Different resolutions in time and frequency are stored, and a subset of raw data is also stored.



SETI@home/Breakthrough!

Searching for Pulses / Triplets
Doppler drift rate -12.0163 Hz/sec Resolution 11,176 Hz
New Pulse: power 2.32, period 11,2146, score 0.52



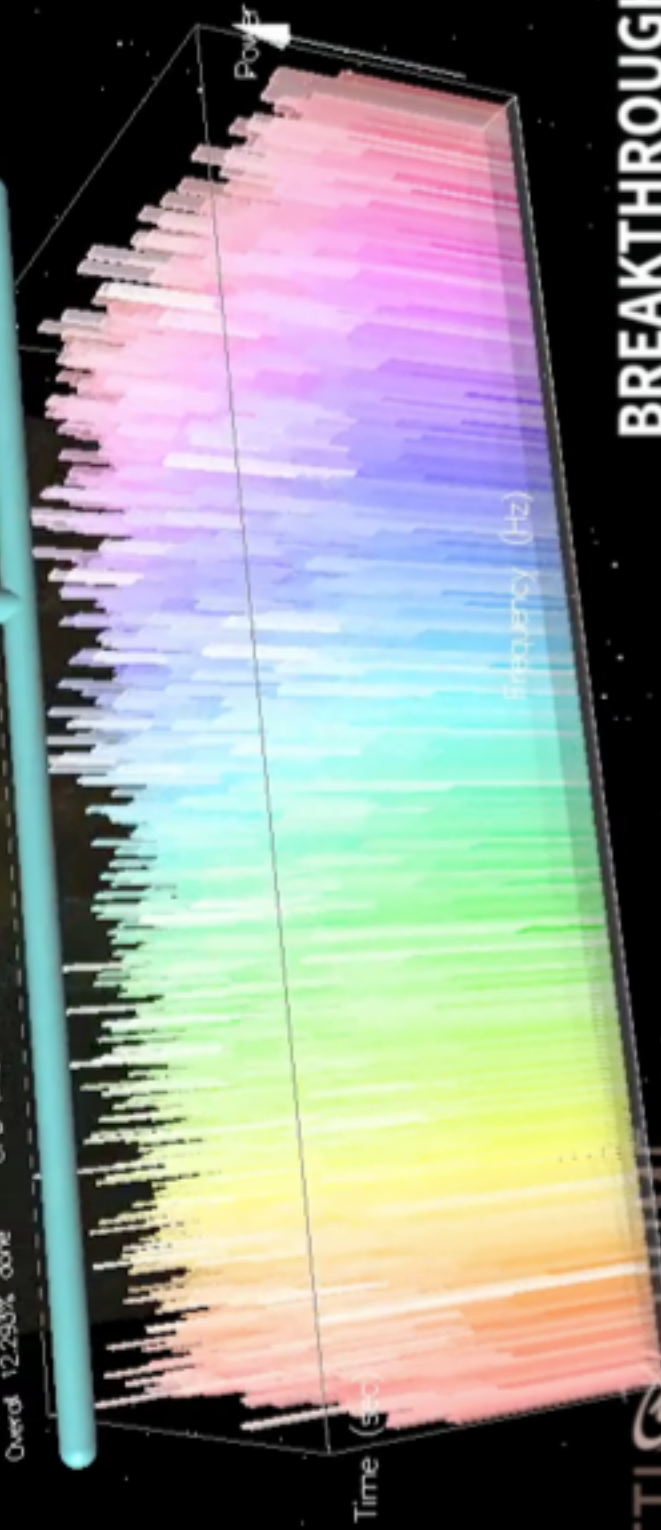
Overall 12.293% done CPU time: 48.92 sec

Data info

From: 17 hr 11' 58" RA, +11 deg 56' 57" Dec
Recorded on: Wed Dec 30 20:45:28 2015
Recorded at: Green Bank Telescope, Rivg, W, Rd 0
Base frequency: 8.568560028 GHz

User info

Name: Eric Korpela
Team: GPU Users Group
Total credit: 1522947564



SETI@HOME

BREAKTHROUGH
LISTEN

BREAKTHROUGH LISTEN DATA RATES: GREEN BANK

RAW DATA	20 GB/SEC
REDUCED SPECTRA	~660 GB/HR
TOTAL ON-SITE STORAGE	2.2 PB

NEED TO MAKE ALL DATA PUBLICLY AVAILABLE!

310
DIE
AAAA

00

8



A

THE PROBLEM WITH FITS



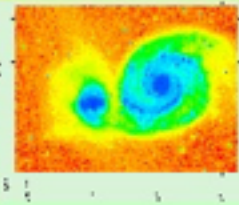
©1979 ATARI INC



FITS

The Astronomical
Image and Table Format

Flexible Image Transport System



- HISTORY FITS was invented in 1979. At the time, there were limitations that
- HISTORY seem a little archaic now. A good example is the FITS header:
- HISTORY KEYWORDS are a maximum of 8 characters and are ALL CAPS.
- HISTORY Each LINE must be shorter than 80 characters.
- HISTORY UNICODE was invented a decade later so isn't supported.
- HISTORY While here have been some minor additions over the years,
- HISTORY such as the CONTINUE and HIERARCH keywords, these are
- HISTORY pretty ugly solutions.

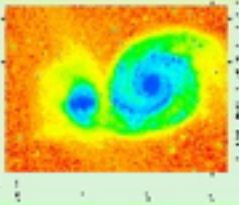


FITS

FITS

The Astronomical
Image and Table Format

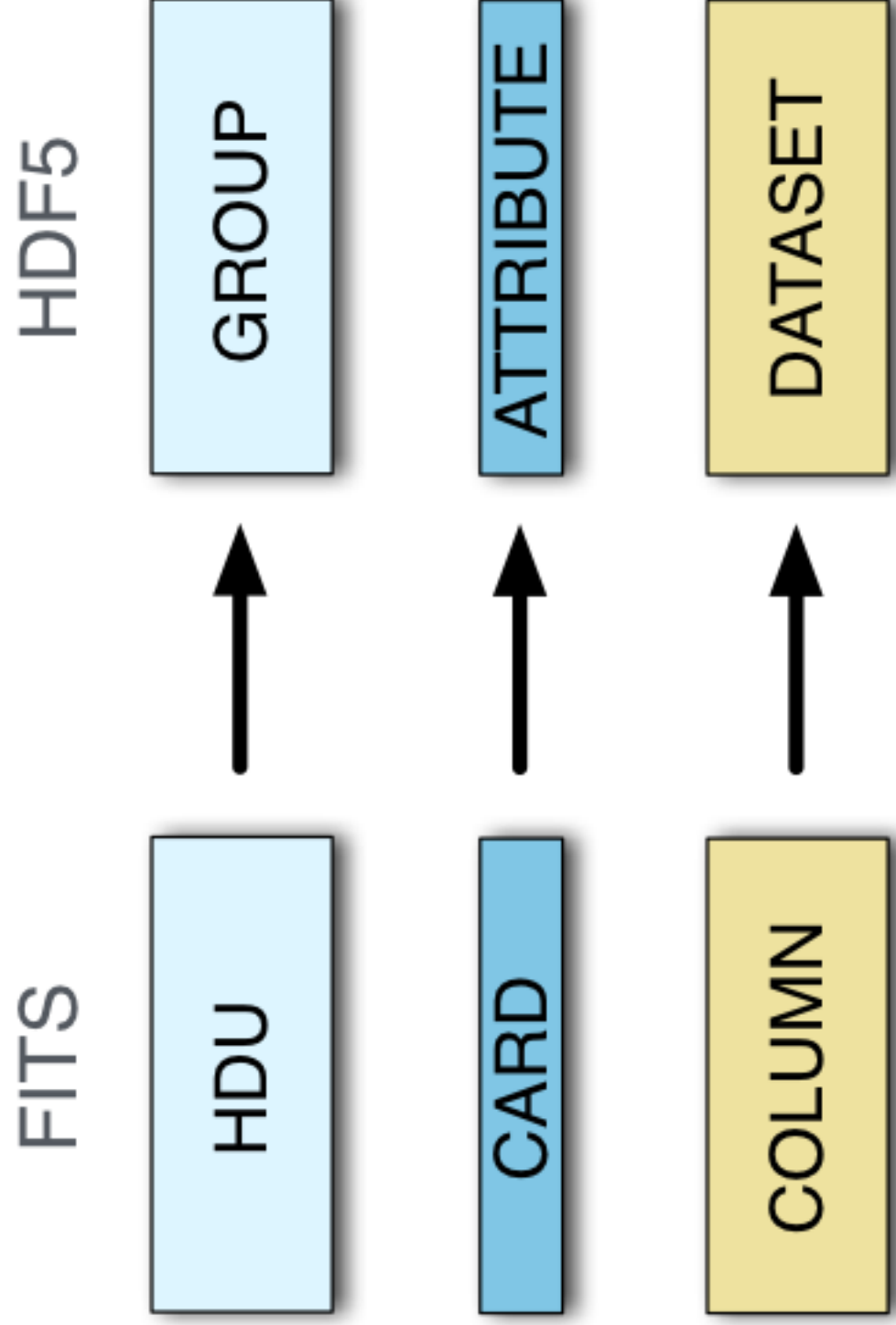
Flexible Image Transport System



- COMMENT For the most part, FITS headers are irritating, not limiting.
- COMMENT A larger problem exists:
- COMMENT Datasets just keep getting bigger.
- COMMENT FITS is not well equipped to handle >1TB datasets.
- COMMENT Pretty damning summary of the issues in [1] and [2]

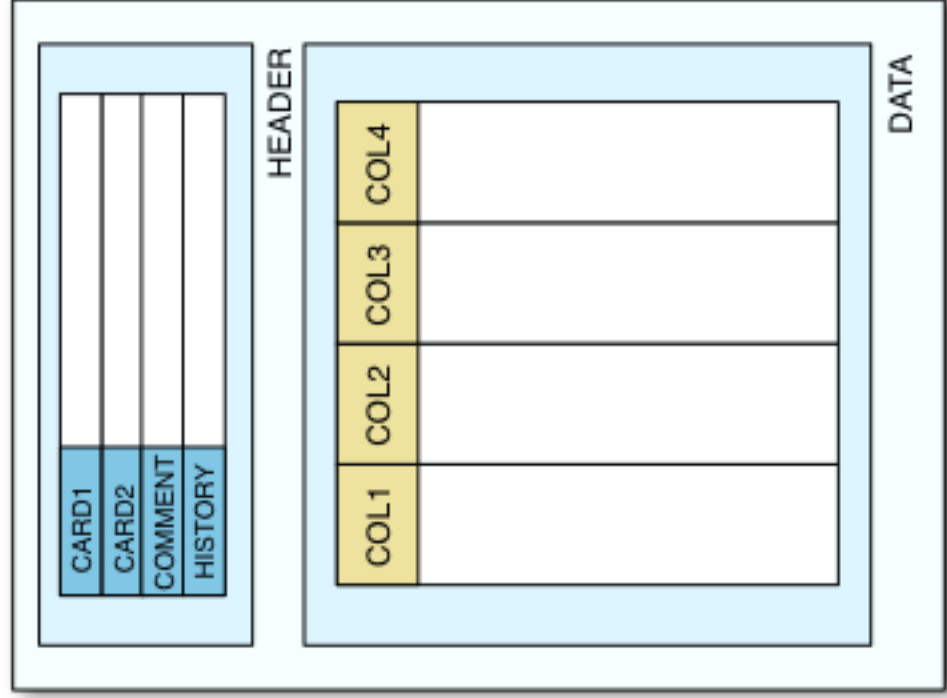
[1] B. Thomas et. al. "Significant Problems in FITS Limit Its Use in Modern Astronomical Research," ADASS XXIII, 2013
[2] B. Thomas and E. al., "The Future of Astronomical Data Formats I. Learning from FITS,"
Astronomy and Computing, pp. 1–14, Apr. 2014.

HDFITS: PORTING FITS DATA MODEL TO HDF5



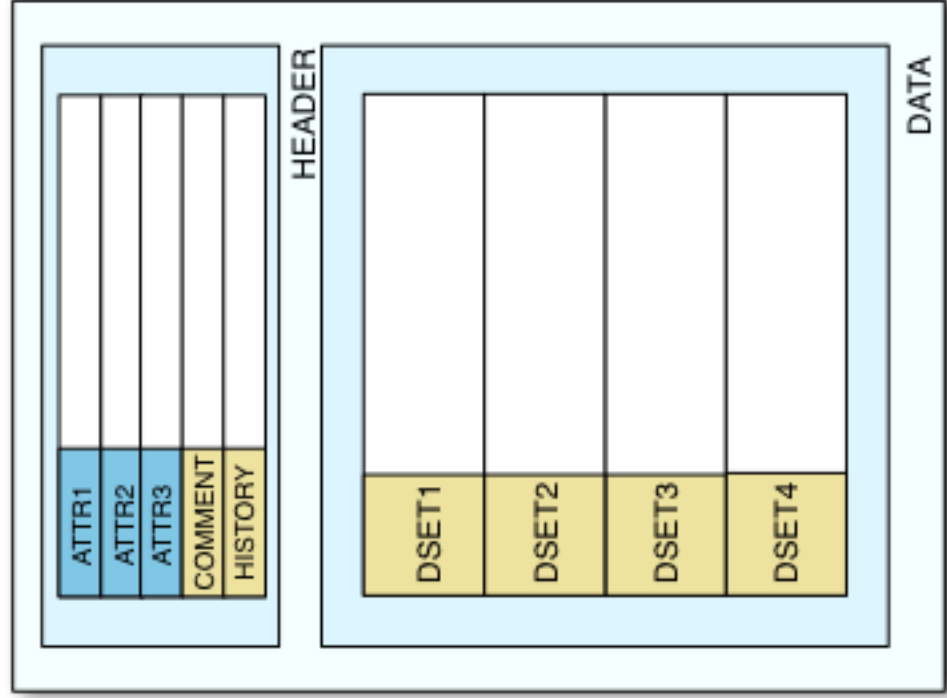
github.com/telegraphic/fits2hdf

FITS



FITS HDU: HEADER DATA UNIT

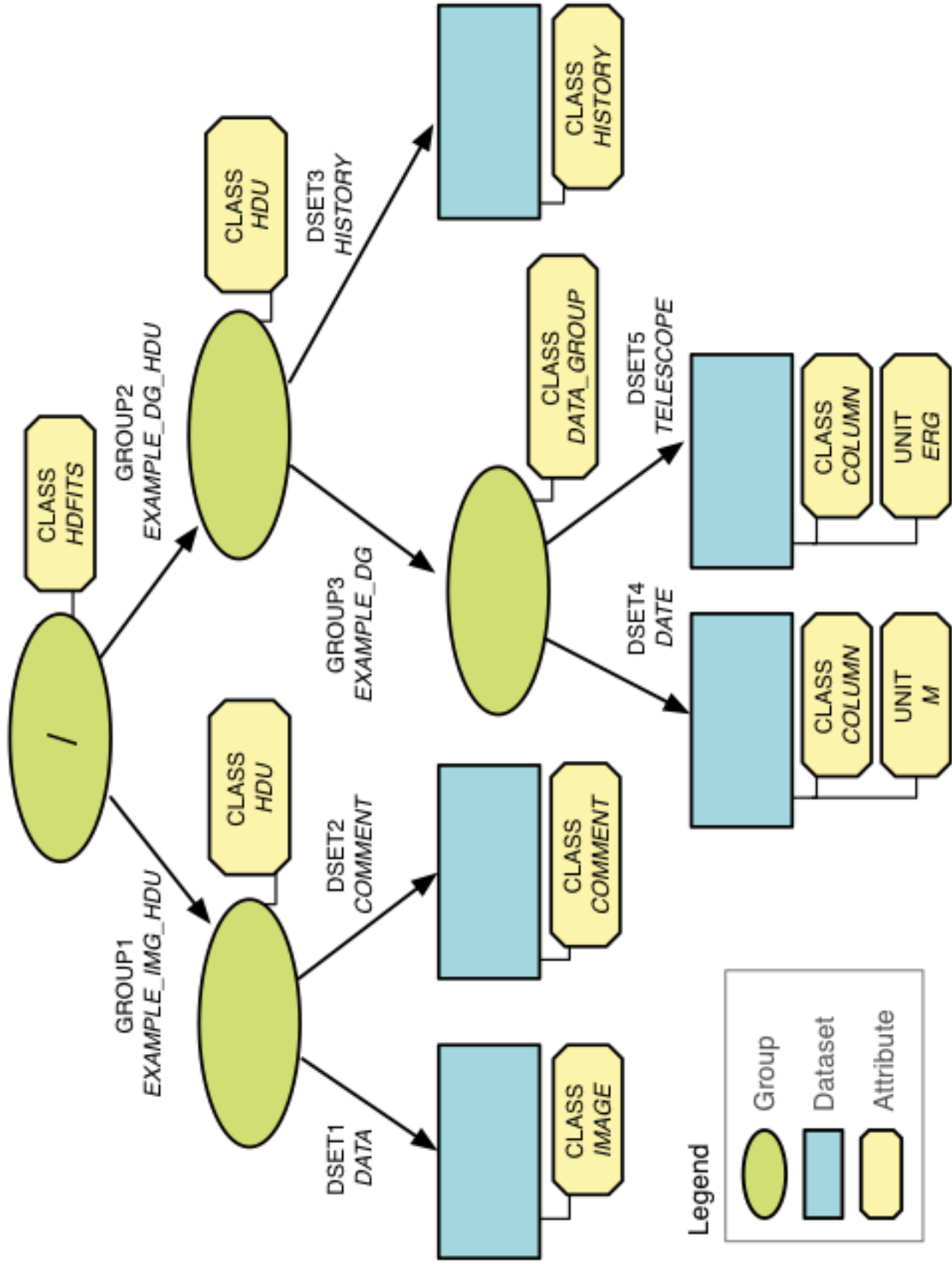
HDF5



HDF5 GROUP: HEADER DATA UNIT

github.com/telegraphic/fits2hdf

HDFITS: EXAMPLE TREE DIAGRAM



CONCLUDING REMARKS

CONCLUDING REMARKS

- Breakthrough Listen is the most comprehensive SETI experiment ever undertaken.
- With big searches comes big data challenges.
- HDF5 will be used instead of FITS!
- Open-source *fits2hdf* conversion utility available at github.com/telegraphic/fits2hdf
- Looking in to OpenDAP + Jupyter in the cloud to provide public data access.



SETI.BERKELEY.EDU | BREAKTHROUGHINITIATIVES.ORG

BACKUP SLIDES

BACKUP SLIDES

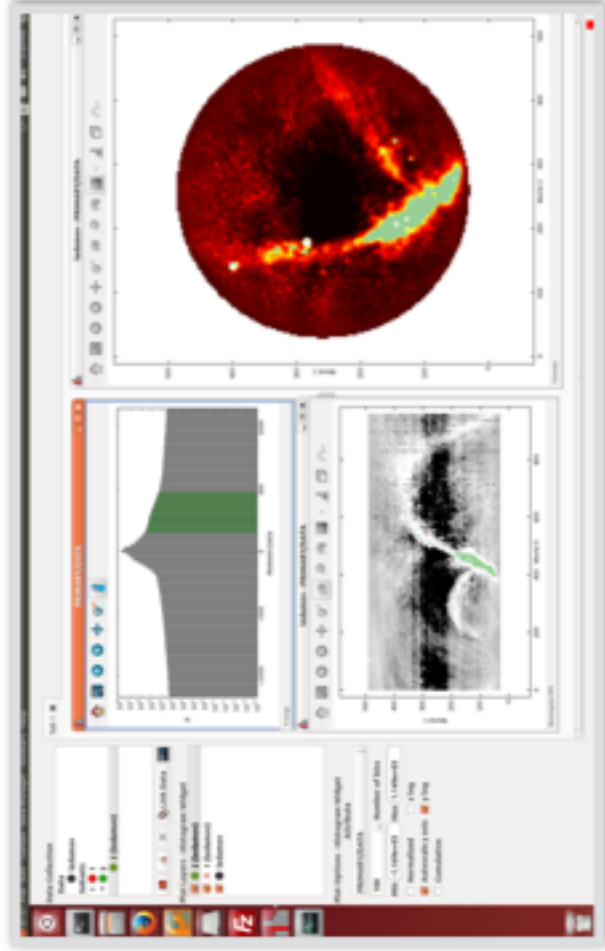
1 in 5 stars harbors an Earth-sized planet in Habitable Zone



I/O Support: Image viewers

I/O support: Image viewers

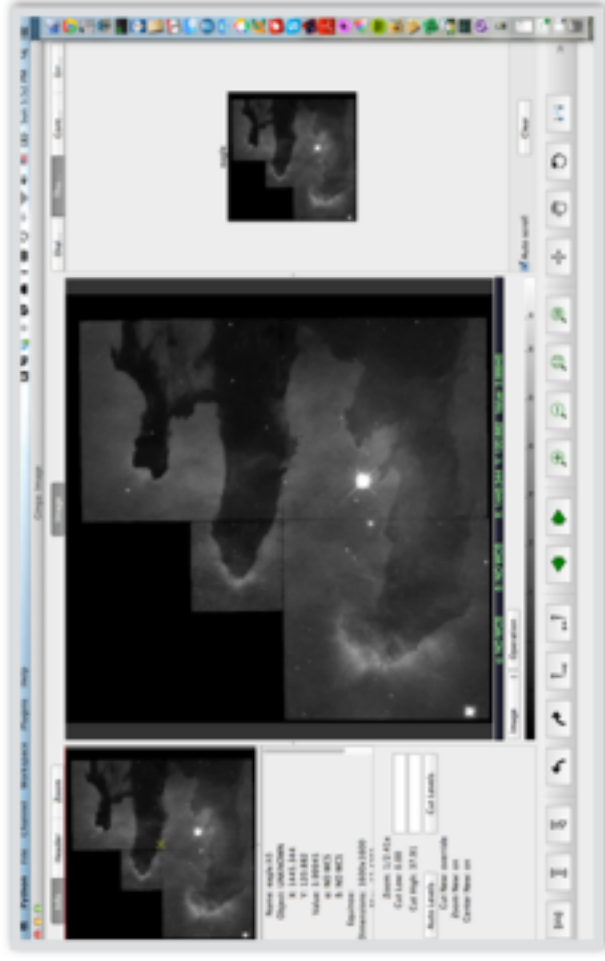
Proof of concept: reading HDFITS images into Python-based FITS viewers



GLUE

www.glueviz.org

2 lines of code changed in order to read HDFITS (thanks to existing basic HDF5 support)



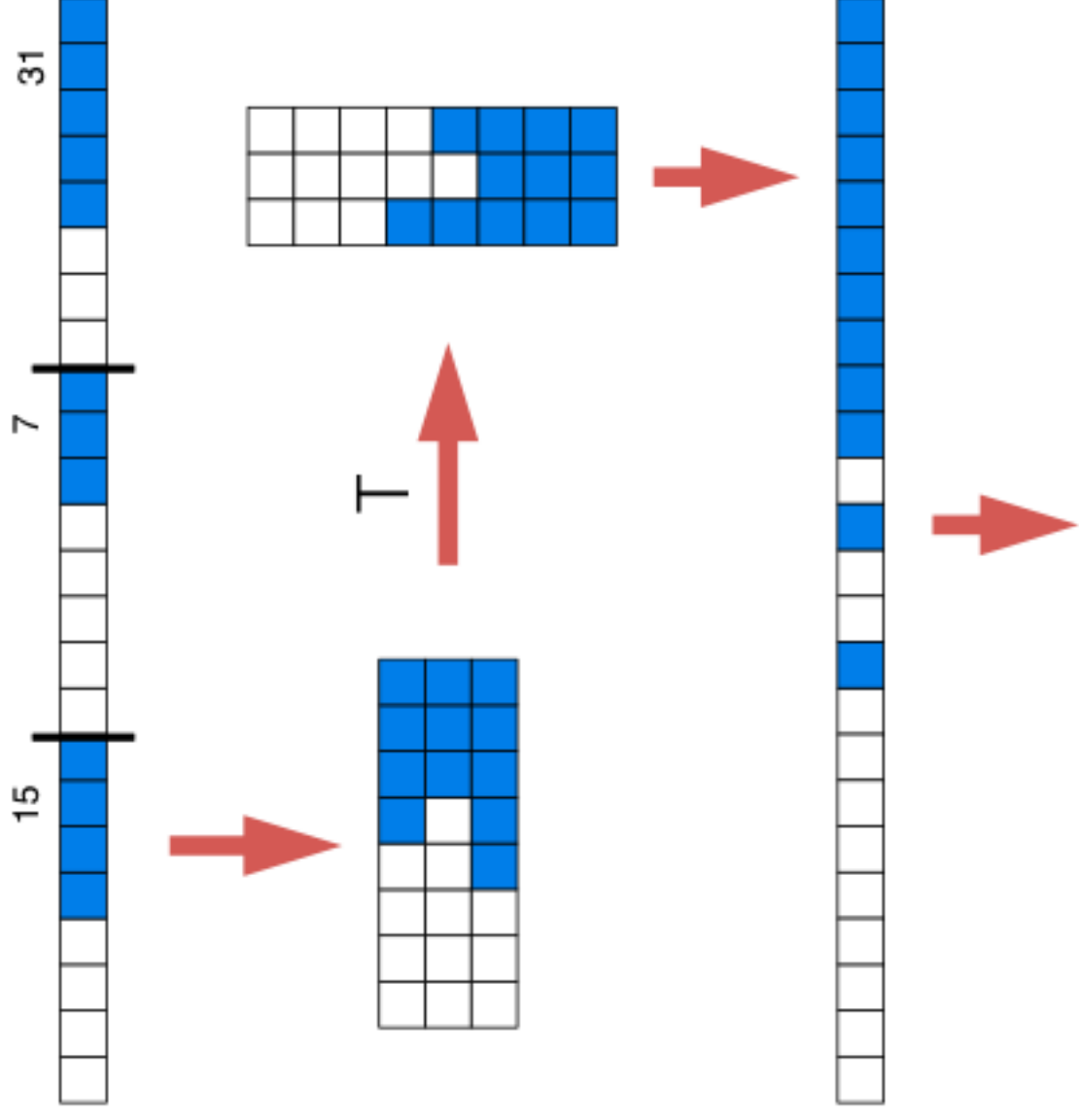
GINGA

ginga.readthedocs.org

10 lines of code changed / added in order to read HDFITS (thanks to abstracted Astrolmage class)

Bitshuffle: a HDF5 pre-compression filter for radio astronomy

for radio astronomy



To compressor

Kiyo Masui (CHIME)
github.com/kiyo-masui/bitshuffle

Results: Compression

Results: Compression

Input file: LEDA visibility data (109 channel, 10 integration subset), in FITS-IDI format

	Size	Ratio	Write time	Read time	Weissman
(FITS)	1.16 GB	1.0 x	8.0 s	7.0 s	-
None	1.16 GB	1.0 x	9.0 s	7.5s	-
LZF	1.08 GB	1.08 x	13.5 s	8.0 s	1.05
bitshuffle	705 MB	1.65 x	7.5 s	9.0 s	2.46
GZIP	831 MB	1.40 x	53.0 s	10.0 s	1.00



Weissman
SCORE™

$$W = \alpha \frac{r \log T}{\bar{r} \log T}$$